

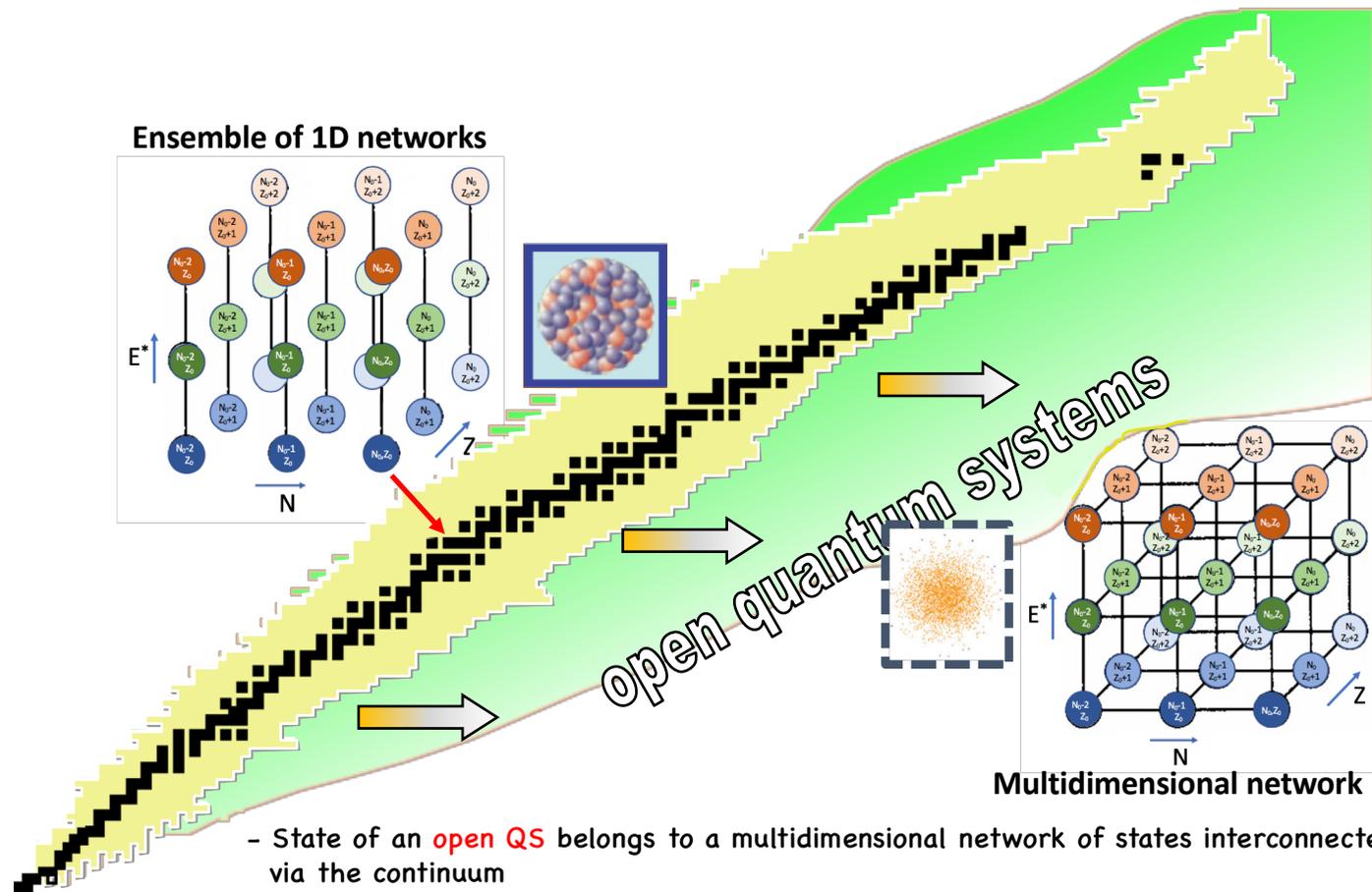
Challenges for nuclear theory

Marek Płoszajczak (GANIL)

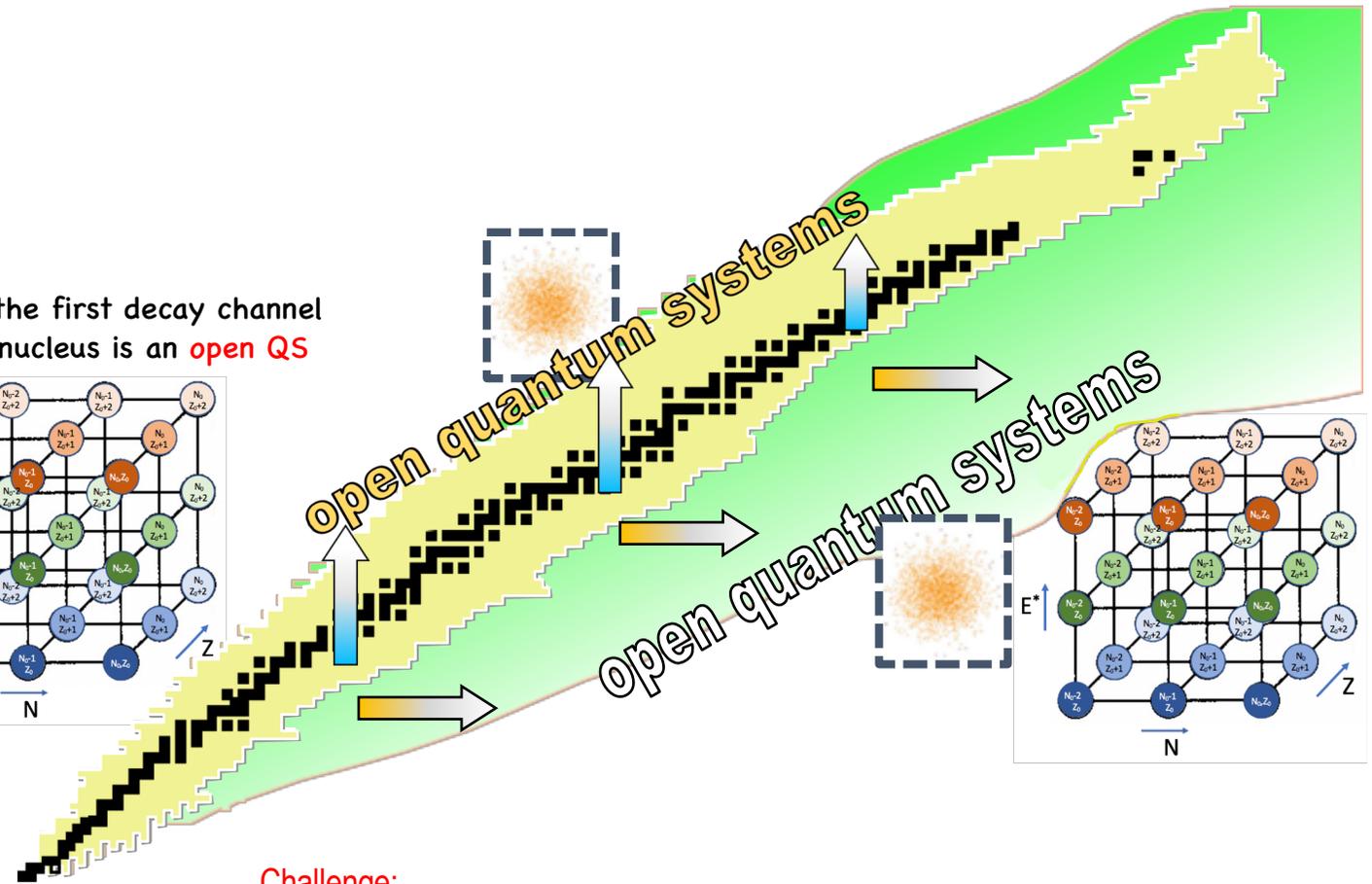
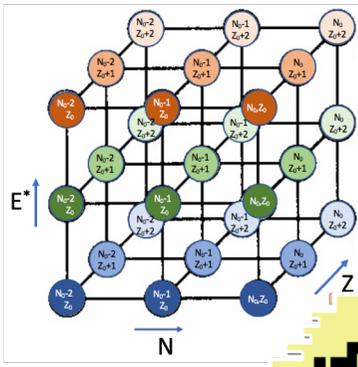
Content:

1. Atomic nucleus: the open quantum system perspective
2. Quantum control of nuclear properties
3. In medium nucleon-nucleon interaction
4. Spectroscopy of super-heavy nuclei
5. Dark matter interactions with nucleons and nuclei
6. Interface between theory and experiment

Atomic nucleus: the open quantum system perspective

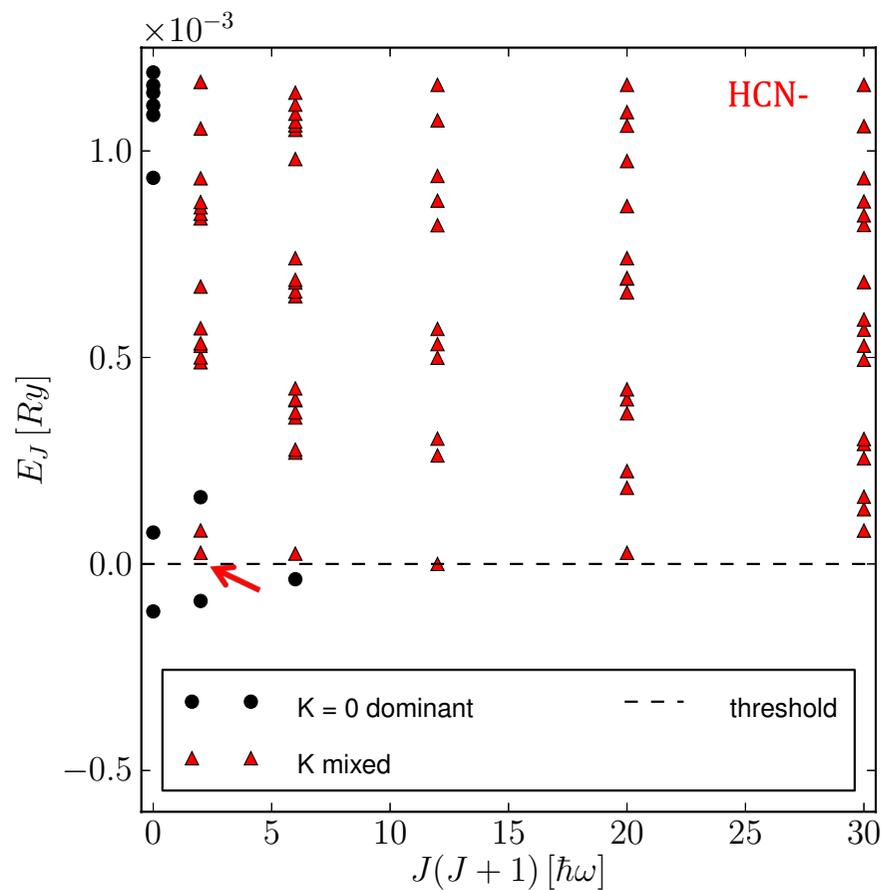


Above the first decay channel
atomic nucleus is an **open QS**



Challenge:
Properties of clusters of correlated states in the multidimensional
network in different domains of E^* , N , Z ?

Bands of resonant states in the dipolar anion



HCN-

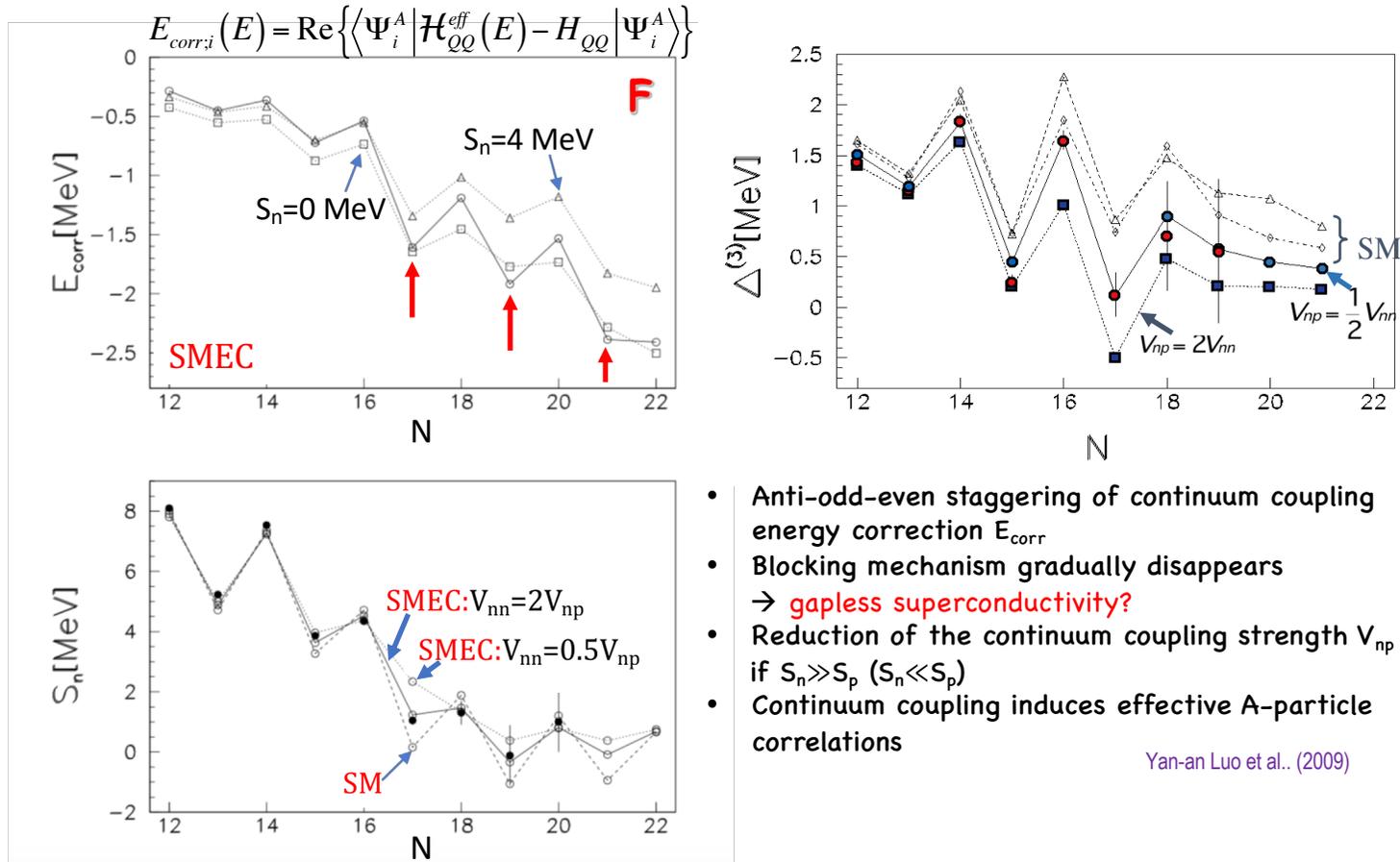
Resonances in rotational bands of dipolar anions are strongly K mixed

Challenge:

Are the γ -selection rules for in- and out- band **nuclear** transitions in the resonance bands unchanged?

What is the nature of near-threshold γ decays?

Continuum coupling correction to binding



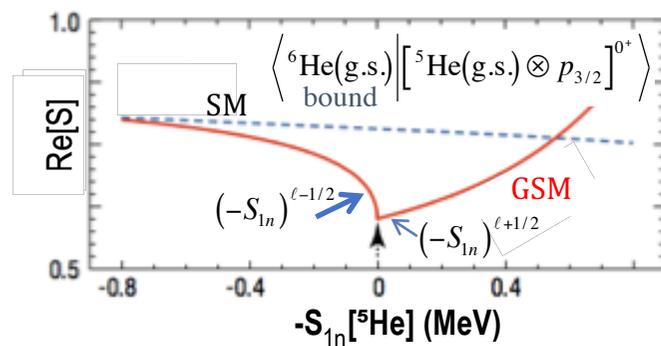
- Anti-odd-even staggering of continuum coupling energy correction E_{corr}
- Blocking mechanism gradually disappears
→ **gapless superconductivity?**
- Reduction of the continuum coupling strength V_{np} if $S_n \gg S_p$ ($S_n \ll S_p$)
- Continuum coupling induces effective A-particle correlations

Yan-an Luo et al.. (2009)

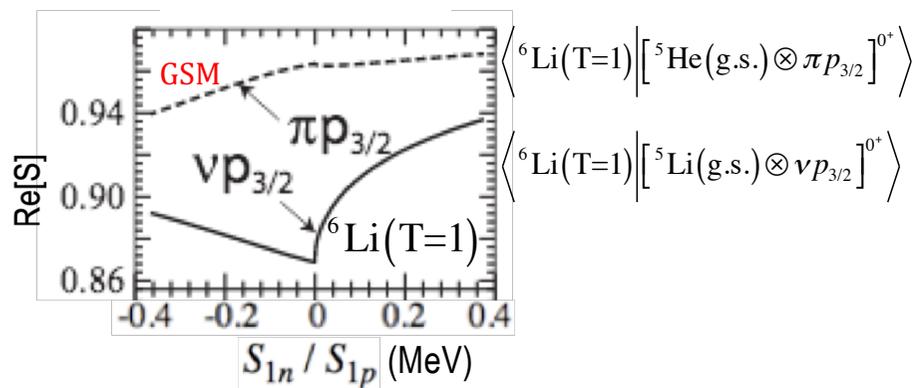
Challenge:

What is the nature of pairing correlations in **open QS**?

Occupation of s.p. shells in near-threshold states



- The **interference phenomenon** between resonant states and non-resonant continuum in the vicinity of the particle emission threshold
- Analogy with the Wigner threshold phenomenon for reaction cross-sections

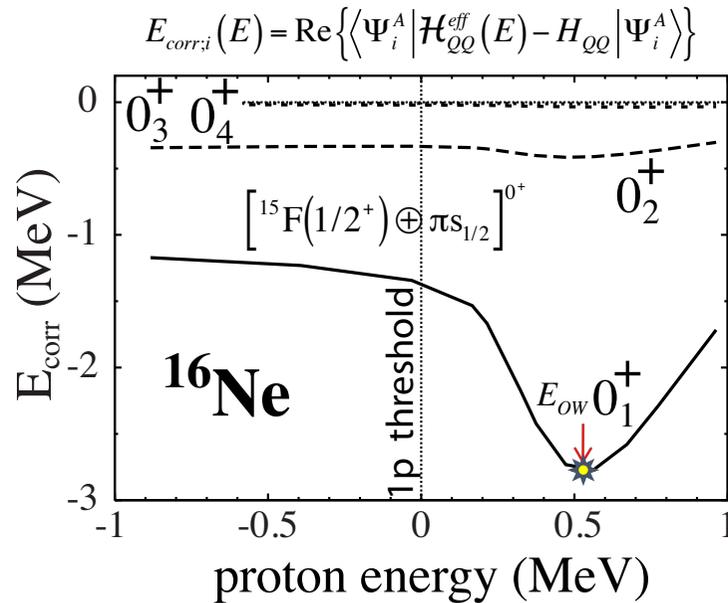


→ Near-threshold configuration mixing acts differently at the proton and neutron drip lines
 N. Michel et al. (2010)

Challenge:

Violation of the mirror symmetry by the coupling to the continuum

Mechanism of the near-threshold collectivization



Okolowicz et al., Prog. Theor. Phys. Suppl. 196 (2012) 230
Fortschr. Phys. 61 (2013) 66

- Interaction through the continuum leads to the formation of the **collective eigenstate** ('aligned state') which couples to the decay channel and carries many of its characteristics
- Point of the **strongest collectivity** (centroid of the 'opportunity energy window') is determined by an interplay between the competing forces of **repulsion** (Coulomb and centrifugal int.) and **attraction** (continuum coupling)
- Emergence of **new energy scales** related to the configuration mixing via decay channel(s)
- This **generic mechanism** explains why so many states both on and off the nucleosynthesis path, exist 'fortuitously' close to open channels

Challenge:

What is the nature of multi-nucleon correlations and clustering in the vicinity of particle emission thresholds?

Quantum control of nuclear properties

- Nuclei, atoms:

- * well defined state
- * limited tunability



- Artificial quantum systems:

(quantum dots, atomic clusters, ...)

- * no "identical" systems
- * wide tunability

- Excitation of highly stripped nuclei in channeling along the crystal axis:

Example: Inverse internal conversion of 73 eV isomeric state ($T_{1/2}=26$ min) of ^{235}U in a laser generated plasma environment

Goldanski, Namiot (1976); Cue, Poizat, Remilleux (1989); Zhu-Shu Yuan, Kimball (1993)

- Application of laser radiation to low-energy nuclear physics:

Matinyan (2018)

- * Laser-induced nuclear anti-Stokes transitions, laser-induced and laser-assisted internal conversion

- * Electron Bridge mechanism:

Krutov, Fomenko (1968); Crasemann (1973)

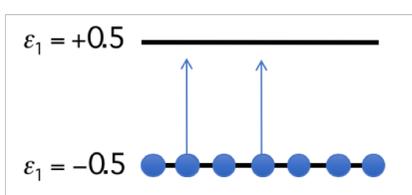
Example: Nuclear isomer excitation in ^{229}Th by intense laser fields

Andreev et al (2018)



Nucleus becomes a distinguished part of a larger system, i.e. nucleus is **the open quantum system** which is found in interaction with an external quantum system, **the environment**

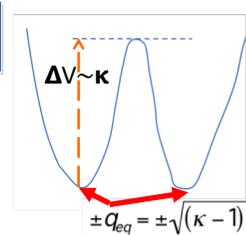
Quantum tunneling in the driven many-body system



$\varepsilon_1 = +0.5$

$\varepsilon_1 = -0.5$

$|\psi_{SD}(z)\rangle \equiv |z\rangle = \exp(z^\dagger G_{21})|0\rangle \Rightarrow \lim_{N \rightarrow \infty} \Psi_{g.s}^{(SU(2))} \equiv \Psi_{HF}^{(SU(2))}$



$\Delta V \sim \kappa$

$\pm q_{eq} = \pm \sqrt{(\kappa - 1) / \kappa}$

\longleftrightarrow

$$\hat{H} = \sum_{k=1}^2 \varepsilon_k \left(\sum_{n=1}^N a_{nk}^\dagger a_{nk} \right) - \frac{1}{2} \sum_{k,l=1}^2 V_{kl} \left(\sum_{n=1}^N a_{nk}^\dagger a_{nl} \right)^2$$

$V_{kl} = V(1 - \delta_{kl}) \quad V \geq 0$

$$G_{kl} = \sum_{n=1}^N a_{nk}^\dagger a_{nl} \Rightarrow \begin{cases} K_0 = \frac{1}{2}(G_{22} - G_{11}) \\ K_+ = G_{21} \\ K_- = G_{12} \end{cases}$$

$\hat{H} = \varepsilon K_0 - \frac{1}{2} V (K_+^2 + K_-^2) \quad \varepsilon = \varepsilon_2 - \varepsilon_1$

$$H_{cl} = \frac{\langle \psi_{SD} | \hat{H} | \psi_{SD} \rangle}{N\varepsilon} = -\frac{1}{2} + \frac{1}{2}(1-\kappa)q^2 + \frac{1}{2}(1+\kappa)p^2 + \frac{1}{4}\kappa(q^4 - p^4)$$

$\kappa = \frac{V(N-1)}{\varepsilon}$

$$V(q) = H_{cl}(q, p=0) = \frac{1}{4}\kappa q^4 + \frac{1}{2}(1-\kappa)q^2 - \frac{1}{2}$$

Time-dependent driving

$$\hat{H}(t) = H\left(t + \frac{2\pi n}{\beta}\right) \quad n = 0, \pm 1, \pm 2, \dots$$

$$H_{cl}(t) = H_{cl}\left(t + \frac{2\pi n}{\beta}\right)$$

$$V(t) = V_0 + \frac{\alpha\varepsilon}{N-1} \sin^2(\beta t) \Leftrightarrow \kappa(t) = \kappa_0 + \alpha \sin^2(\beta t)$$

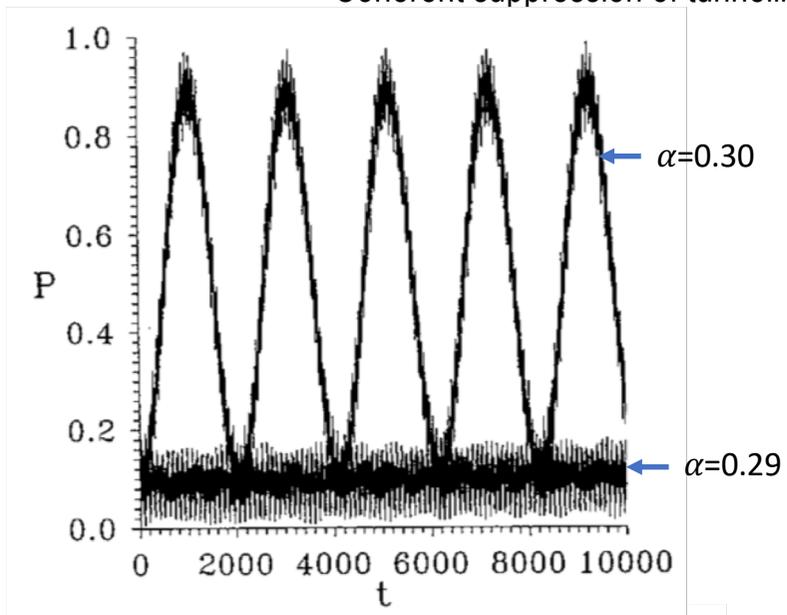
$$\Psi(t) = \sum_k c_k \Omega_k(t) = \sum_k c_k \exp(-ie_k t) \Phi_k(t)$$

$\Phi_k(t) = \Phi_k\left(t + \frac{2\pi n}{\beta}\right) \quad n = 0, \pm 1, \pm 2, \dots$

quasi-energies
quasi-energy eigenstates

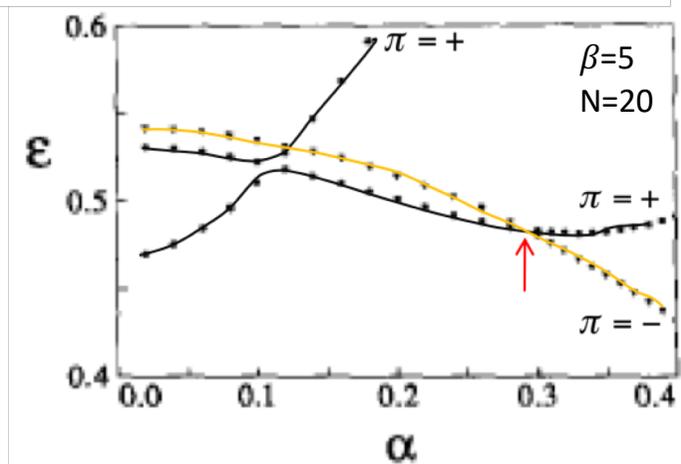
Evolution of the wave packet depends on the energy spectrum of Floquet states which depends on the external environment and **can be tuned**

Coherent suppression of tunneling



By a suitable choice of the driving amplitude one can extinguish the dominant oscillation frequency

Kaminski, MP, Arvieu (1994)



Challenge:

How to engineer the environment of an atomic nucleus in order to control its lifetime?, i.e. can one design an **artificial** atomic nucleus?

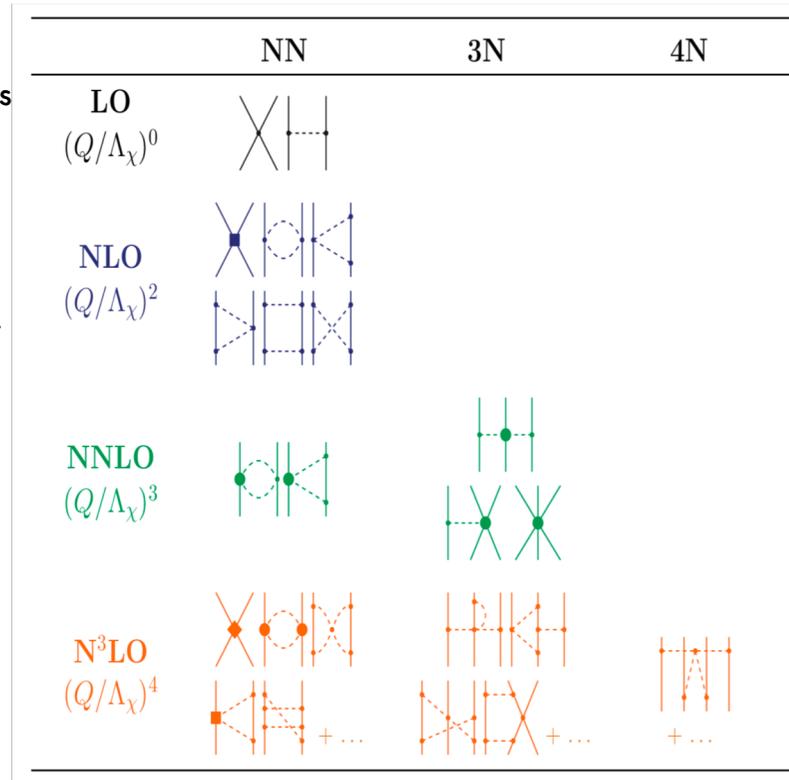
In medium nucleon-nucleon interaction

- First principles for nuclear physics: **QCD**
 - non-perturbative at low energies
 - lattice QCD in the future; first applications to spectrum, magnetic moments, polarizabilities ($A < 5$) and $np \rightarrow d\gamma$, β -decay of ${}^3\text{H}$, ... for unphysical quark masses: $m_\pi \sim 450 \text{ MeV}$, $m_N \sim 1200 \text{ MeV}$

- For now inter-nucleon forces from chiral effective field theory (χEFT)
 - based on the symmetries of QCD; χ -symmetry of QCD ($m_u \approx m_d \approx 0$) broken with pion as a Goldstone boson
 - degrees of freedom: nucleons and pions

but

- χEFT is not regularizable; dependence on cutoffs
- Systematic low-momentum expansion to a given order (Q/Λ_χ) is still debated
 - ➔ the hierarchy of many-body interactions and consistency of the χEFT framework is not yet proven



Challenge:

The determination of EFT scheme which allows for a reliable and systematic low-momentum expansion and defines the hierarchy of many-body interactions

Dispersive optical model perspective on in-medium NN interaction

How neutron (proton) interact in nuclear medium? \longrightarrow (N-Z)-dependence of the optical potential

$$W^{(p)} = \frac{\hbar v \rho_N}{2} \left(\frac{N}{A} \langle \sigma_{np} \rangle + \frac{Z}{A} \langle \sigma_{pp} \rangle \right) = \frac{\hbar v \rho_N}{4} [(\langle \sigma_{np} \rangle + \langle \sigma_{pp} \rangle)] + \frac{N-Z}{A} (\langle \sigma_{np} \rangle - \langle \sigma_{pp} \rangle) = W^{(0)} + \frac{N-Z}{A} W^{(1)}$$

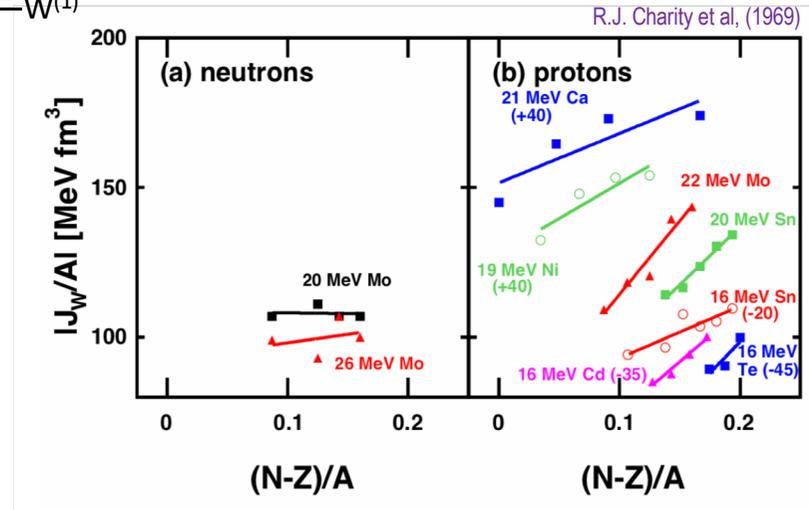
G.R. Satchler (1969)

$$W^{(n)} = \frac{\hbar v \rho_N}{2} \left(\frac{N}{A} \langle \sigma_{nn} \rangle + \frac{Z}{A} \langle \sigma_{np} \rangle \right) = \frac{\hbar v \rho_N}{4} [(\langle \sigma_{np} \rangle + \langle \sigma_{nn} \rangle)] + \frac{N-Z}{A} (\langle \sigma_{np} \rangle - \langle \sigma_{nn} \rangle) = W^{(0)} - \frac{N-Z}{A} W^{(1)}$$

Average int. of projectile nucleon
with individual target nucleons : $W \simeq W^{(0)} \pm \frac{N-Z}{A} W^{(1)}$

Most global optical potentials do not include
the asymmetry dependence of the central
potential

π n ℓ j	⁴⁰ Ca	⁴⁸ Ca
1s _{1/2}	0.73	0.63
0d _{3/2}	0.76	0.69
0f _{7/2}	0.73	0.63
ν n ℓ j		
1s _{1/2}	0.76	0.80
0d _{3/2}	0.78	0.77
0f _{7/2}	0.71	0.80



Protons (neutrons) experience stronger (weaker) correlations in neutron-rich matter

R.J. Charity, W. Dickhoff, (1969)

Challenge:

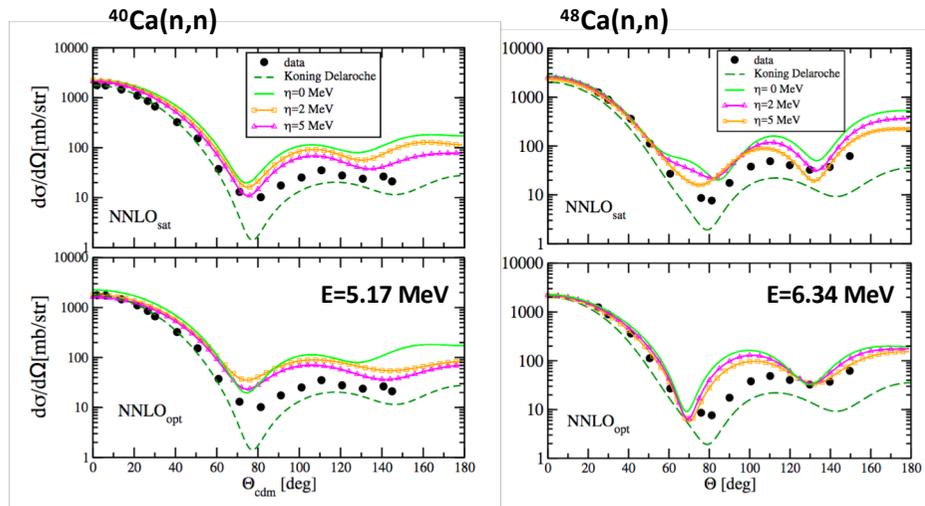
Simultaneous study of neutron/proton elastic cross section, spectral functions and charge distributions in **long isotopic chains**

Optical model from first principles

- New Hamiltonian approaches for nuclear structure and reactions: NCGSM, NCSMC, CC, IM-SRG, ...

→ unification of nuclear structure and reactions in low-energy continuum

Optical potential from CCSD approach



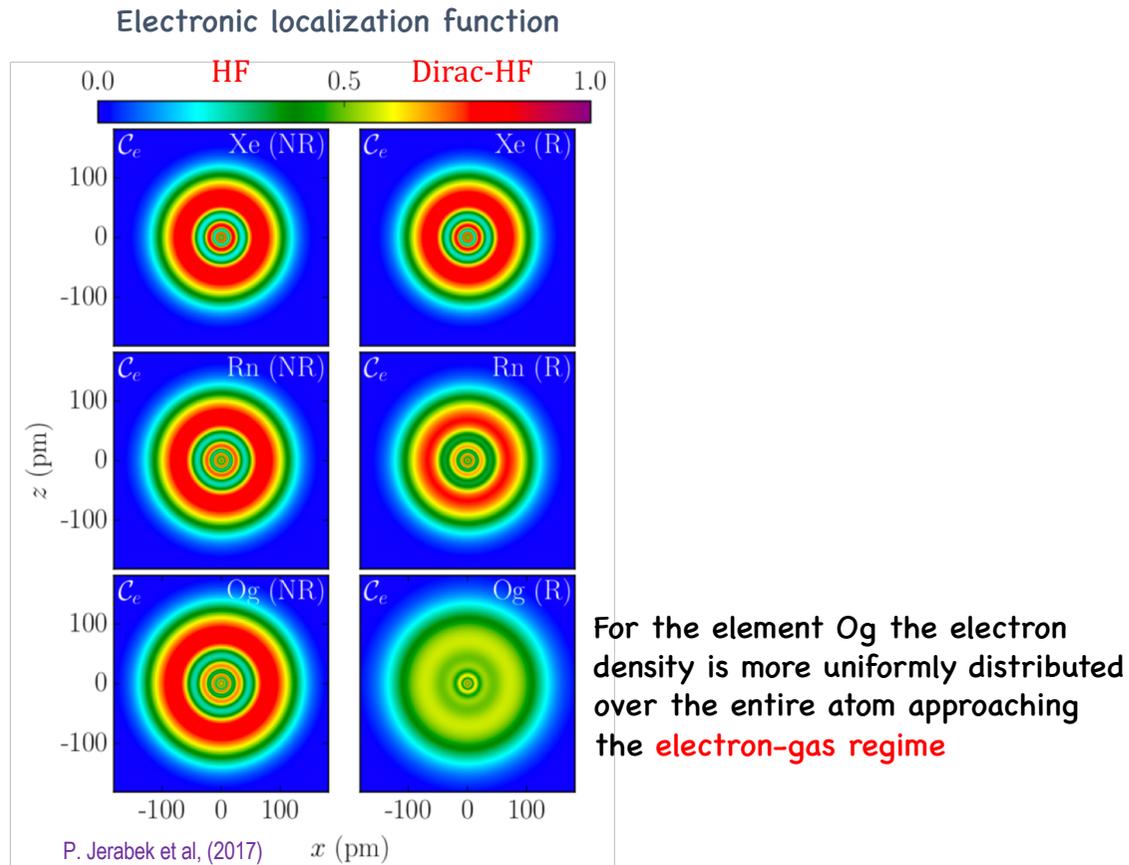
J. Rotureau et al. (2018)

Challenge:

Construction of the microscopic optical potential with the correct absorption properties

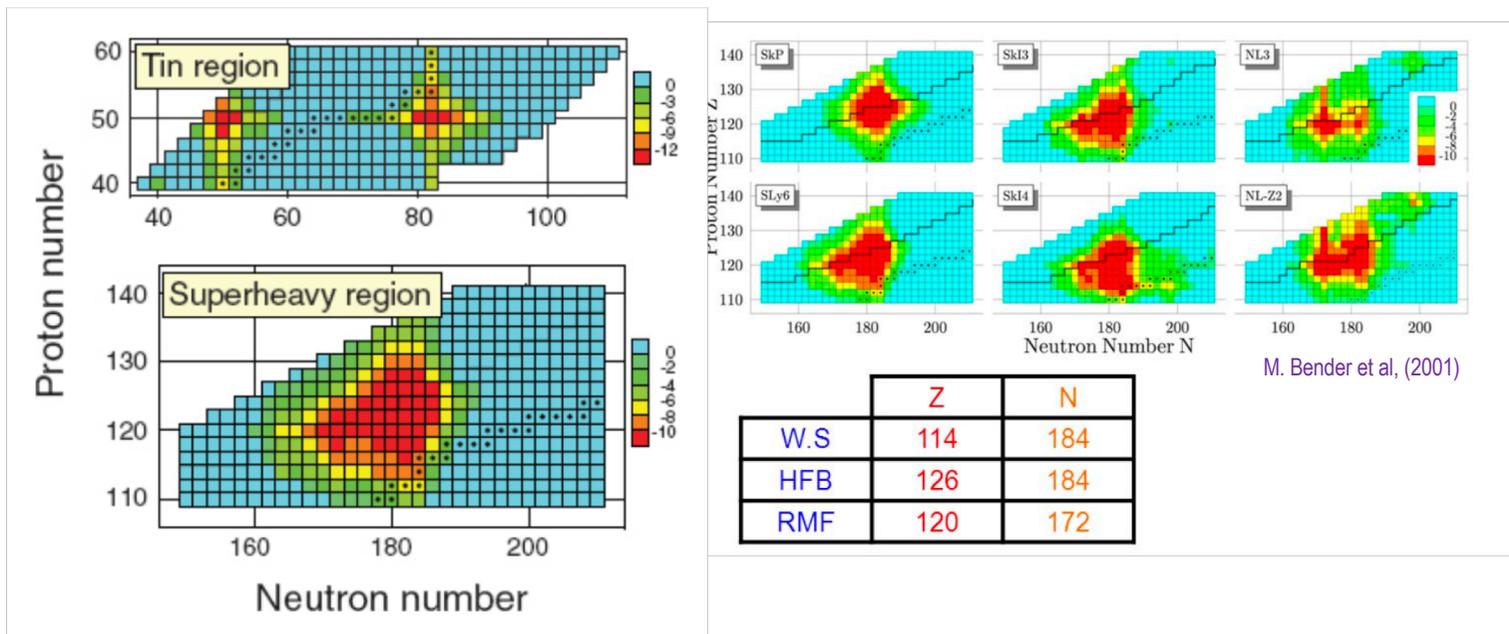
Spectroscopy of superheavy nuclei

- Relativistic effects in super-heavy atoms may change the electronic structure
 - transition from the LS-coupling to jj-coupling of the Dirac equation at large Z
 - large spin-orbit splitting smears out the single-particle density



Is it the end of Mendeleev periodicity law as we know it?

- Single-particle level density grows faster than $A^{1/3}$
 - great sensitivity to the interplay between short-range (**attractive**) nuclear interaction and long-range (**repulsive**) Coulomb interaction
- Disappearance of spin-orbit gaps (magic numbers)
 - great sensitivity to the details of in-medium NN interaction and continuum coupling
 - gradual transition to the uniform-gas regime for nucleonic localization
 - familiar **pattern of shells separated by magic gaps is gone!**

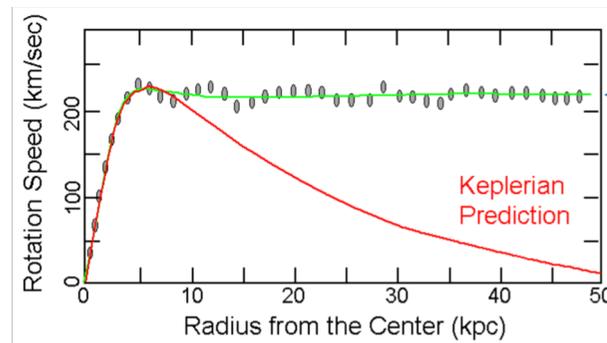


Challenge:

What should be the theoretical approach to analyze and systematize the spectroscopic data in super-heavy mass region?

Dark matter interaction with nucleons and nuclei

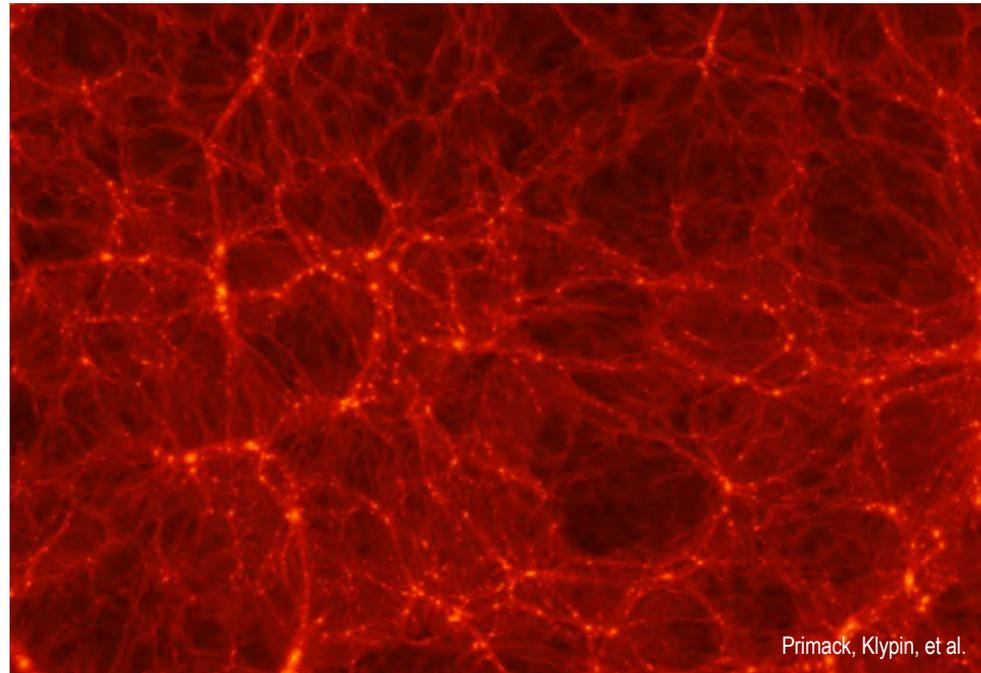
- Perhaps the most-likely-to-be resolved **new-physics** problem
- Closely linked to laboratory-based accelerator and underground experiments
- Existence of dark matter deduced from its dynamical effects in astrophysics:
 - flat velocity rotation curves in galaxies



$V \propto \text{const}$
 $m(r) \propto r$
 $\rho(r) \propto 1/r^2$

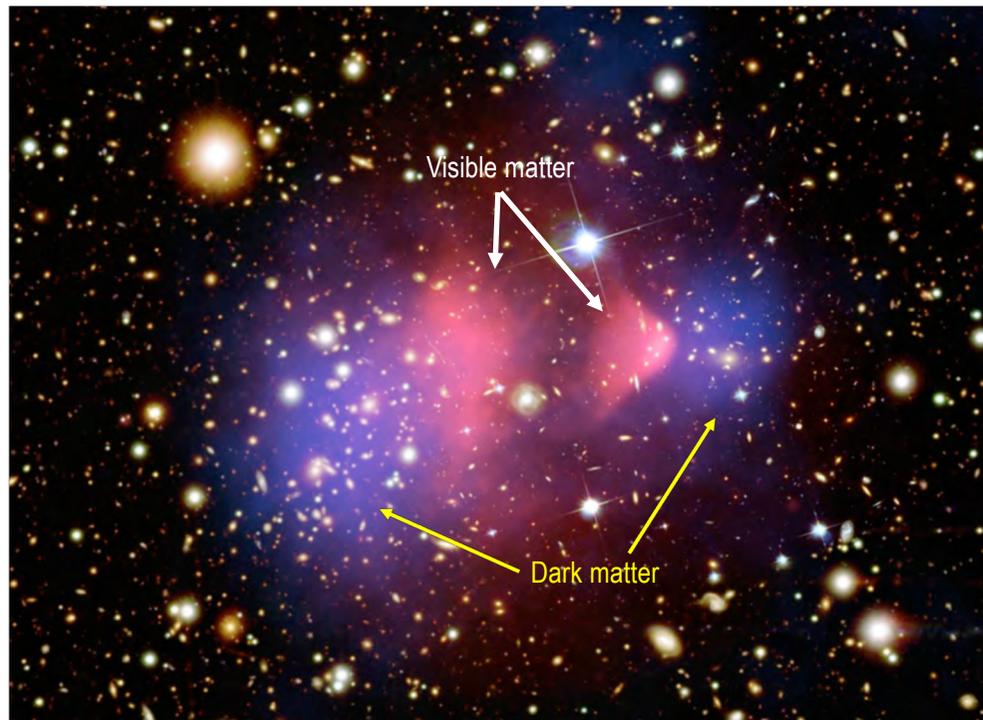
W. Haxton (2014)

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Primack, Klypin, et al.

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 - seen in the difference between gravitating (lensing) and radiating matter distributions in collision of galaxy clusters



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 - seen in the difference between gravitating (lensing) and radiating matter distributions in collision of galaxy clusters
- Bulk of the dark matter must reside beyond the standard model (SM)
- Properties:
 - long-lived or stable
 - cold or warm (slow enough to seed the structure formation)
 - gravitationally active
 - lacks strong couplings to itself or to baryons
 - leading candidates: weakly interacting massive particles (WIMPs) and axions
- Detection:
 - collider searches : SM particles \rightarrow WIMPs
 - astrophysical signals (indirect) : WIMPs \rightarrow SM particles
 - * claims of a dark matter annihilation signal at the galactic center, consistent with a $\sim 30\text{--}40$ GeV WIMP annihilating to b quarks, producing ~ 5 GeV photons
 - nuclear searches : WIMPs + Nucleus(m_A, J, T) \rightarrow WIMPs + Nucleus(recoil) (direct)
 - * world-wide effort to search for WIMPs in elastic scattering on nuclei with different spins, isospins, masses

Challenge:

The determination of WIMP-nucleus interaction

Interface between theory and experiment

- In many cases, nuclear input MUST involve massive extrapolations based on predicted quantities ... and extrapolations are ~~impossible~~ tough.
- Naïve approach to a model error estimate:
 - Take a set of reasonable models M_i
 - Make a prediction $E(y;M_i)$
 - Compute average and variation within this set
 - Compute rms deviation from existing experimental data. If the number of fit-observables is large, statistical error is small and the error is predominantly systematic.
- Bayesian Model Averaging (BMA): what is required?
 - Common dataset (as large as possible) needs to be defined
 - Make a statistical analysis for individual models
 - Make individual model predictions, including statistical uncertainties
 - Decision should be made on the prior model probability (model selection)
 - Model averaging refers to the process of estimating some quantity under each model and then averaging the estimates according to how likely each model is

→ There is no need to choose one model; it is possible to average the predictions from several models

BMA could be very useful to prepare experiments in unknown regions... but **should not be used as an excuse for poor science**

Challenge:

Bayesian neural networks are prone to instabilities and can be very time-consuming

Message to take

- **Theory and experiment are intertwined**

- Theory gives the mathematical formulation of our understanding and predictive ability while experiment forces us to create new theories and provides verification to the existing ones

- **Crucial advances in low-energy nuclear theory have been made in this century:**

- In medium nucleon-nucleon interaction from basic principles (EFT)
 - *Ab initio* structure and reactions (GFMC, NCSM(C), NCGSM, CCM,...)
 - Continuum Shell Model (GSM, SMEC); structure and reactions in the low-energy continuum
 - Many-body description of the large amplitude collective motion (e.g. fusion and fission)

- Tremendous efforts are needed before these conceptual advances will bring a real progress in the quantitative description of atomic nuclei

- **Interaction with other fields of Physics and discoveries/technological progress in nuclear studies raise new challenges, change paradigms and methodology in nuclear studies**

- "No man is an island entire of itself; every man is a piece of the continent, a part of the main..."

J. Donne (1572-1631), MEDITATION XVII, *Devotions upon Emergent Occasions*

- **Costly nuclear-physics projects will compete with other large-scale projects in Science**

- Nuclear physics must demonstrate that it greatly contributes to the global progress of Science and advance societal applications

All of this requires excellent theory!